

# The Origin, Age, and Stratigraphy of Mars' South Polar Cap

Peter Buhler<sup>1</sup>

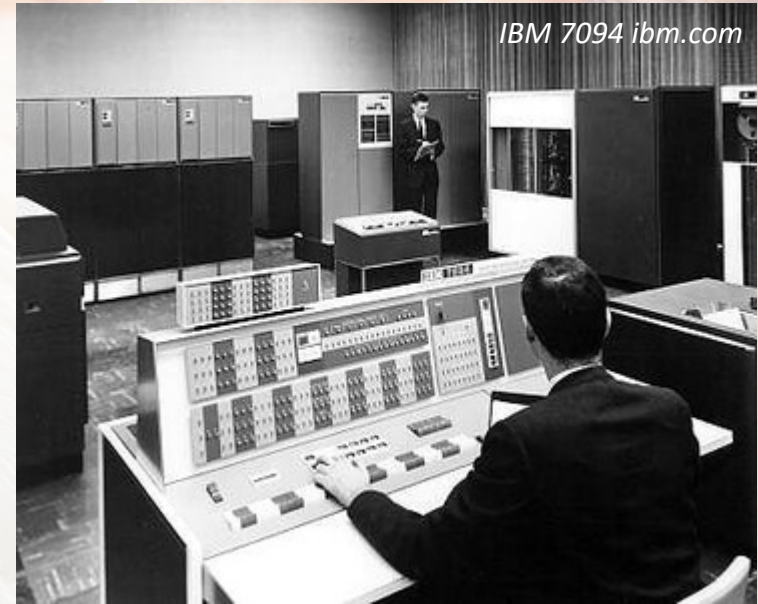
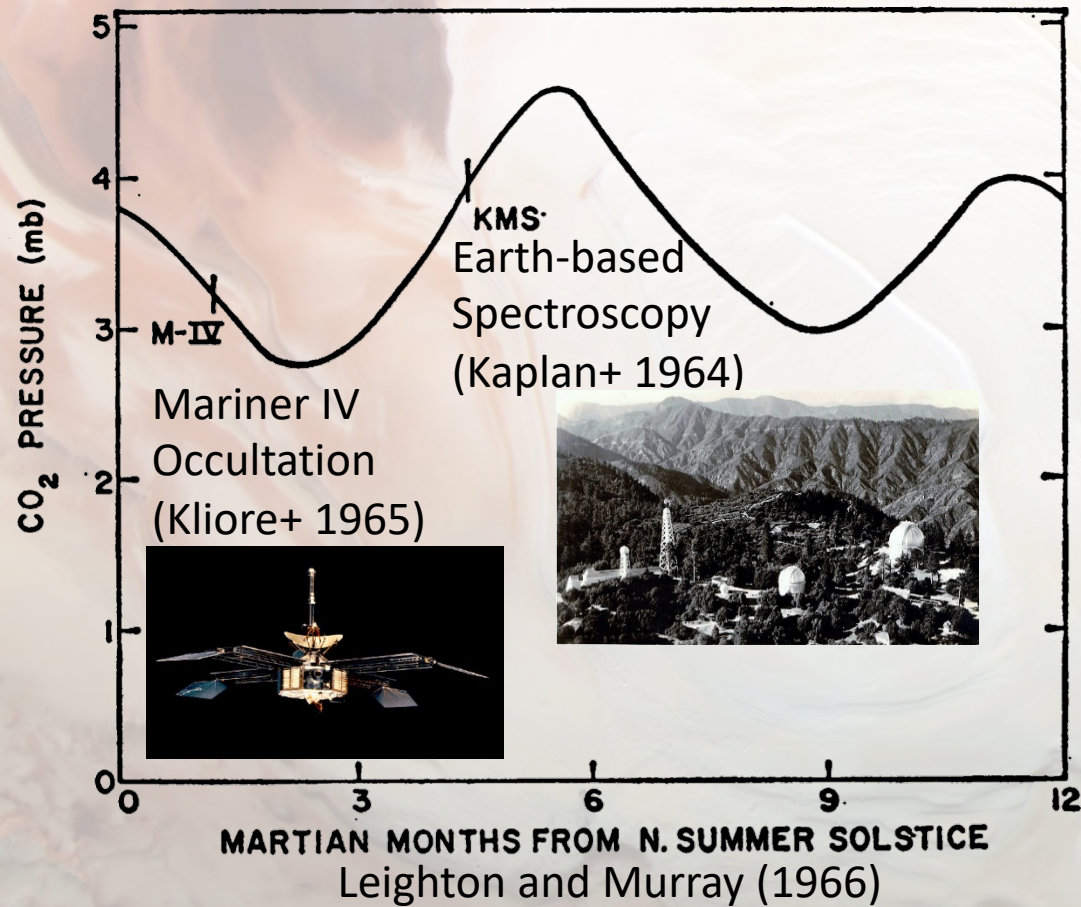
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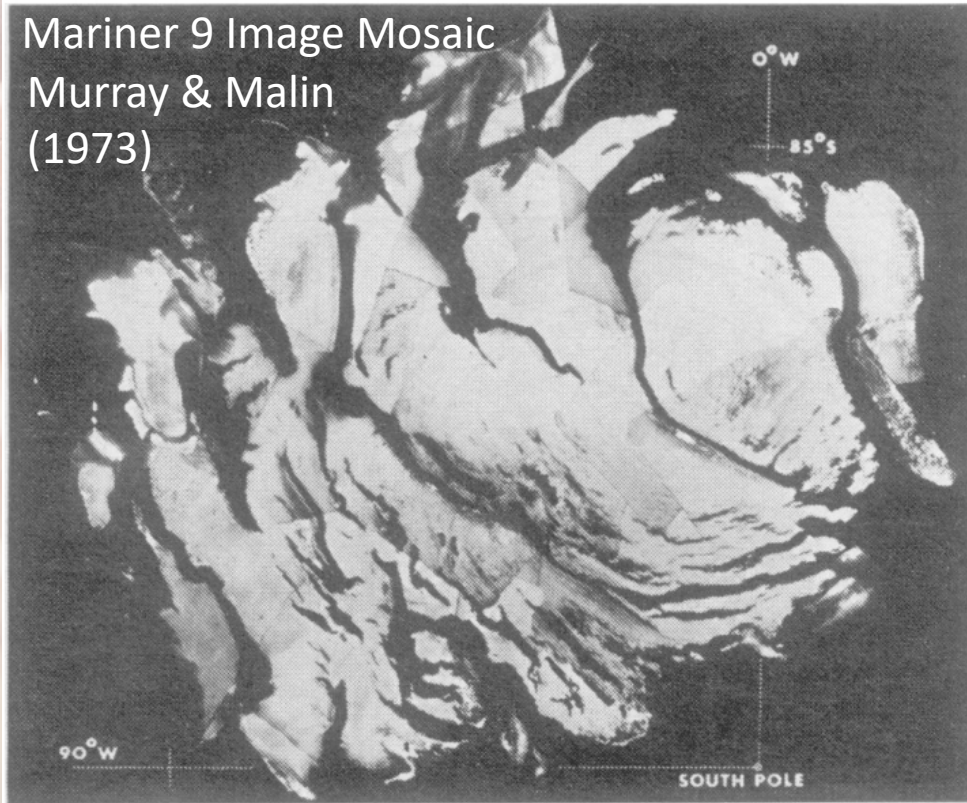
# Background



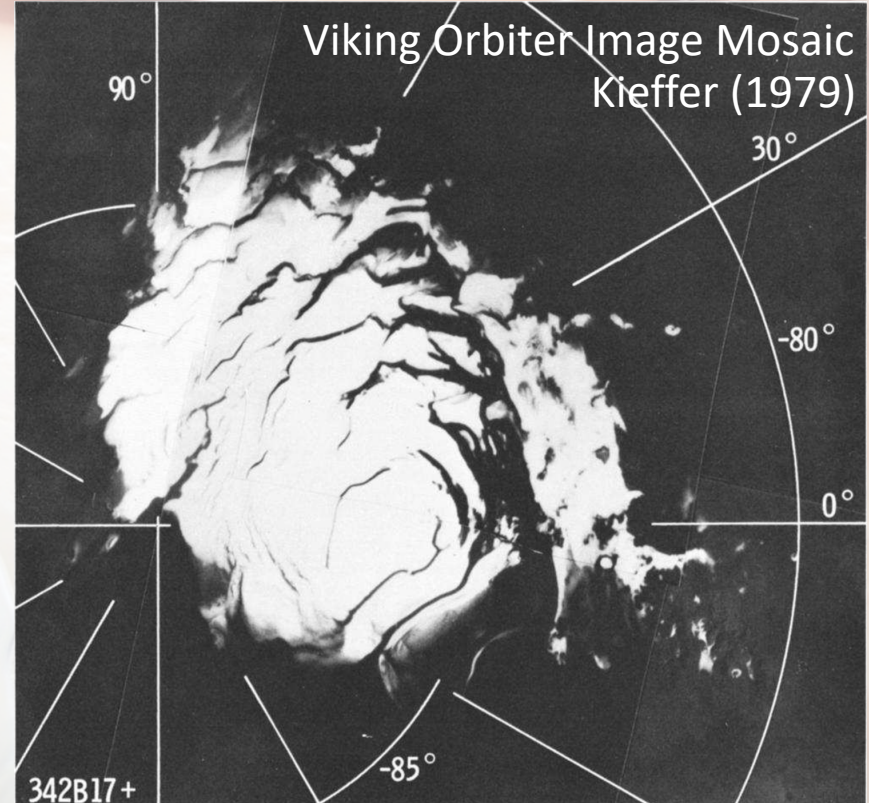


# What is the RSPC?

Mariner 9 Image Mosaic  
Murray & Malin  
(1973)



Viking Orbiter Image Mosaic  
Kieffer (1979)

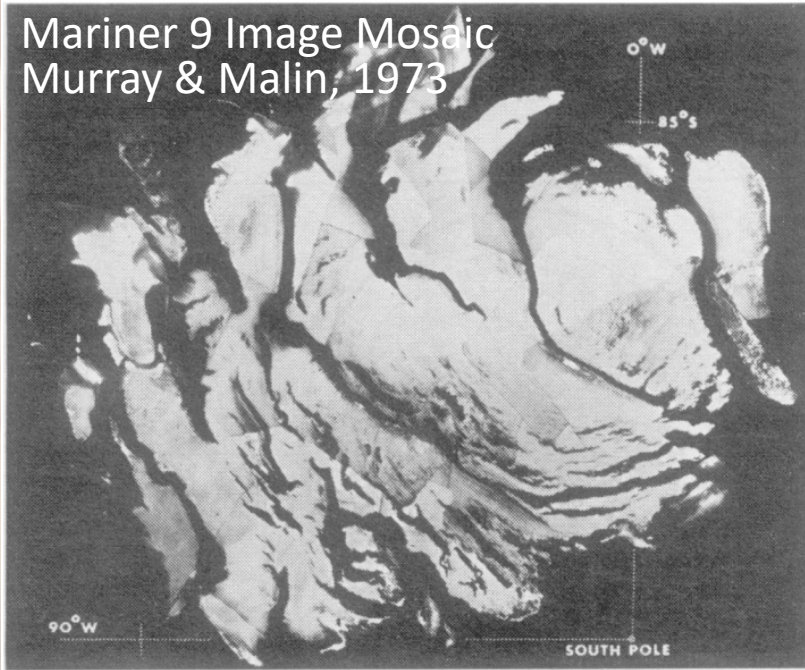


- “CO<sub>2</sub> ice could not survive in contact with low-albedo material.”
- “Therefore, a residual water-ice cap is much more stable than a solid CO<sub>2</sub> one on Mars in the summertime.”
- “Throughout the summer, the polar frost remained at the temperature of solid CO<sub>2</sub>.”
- “Thus Mars appears to have a residual polar cap of CO<sub>2</sub> in the south and one of H<sub>2</sub>O in the north.”



# OK, it's CO<sub>2</sub>, but why at the South Pole?

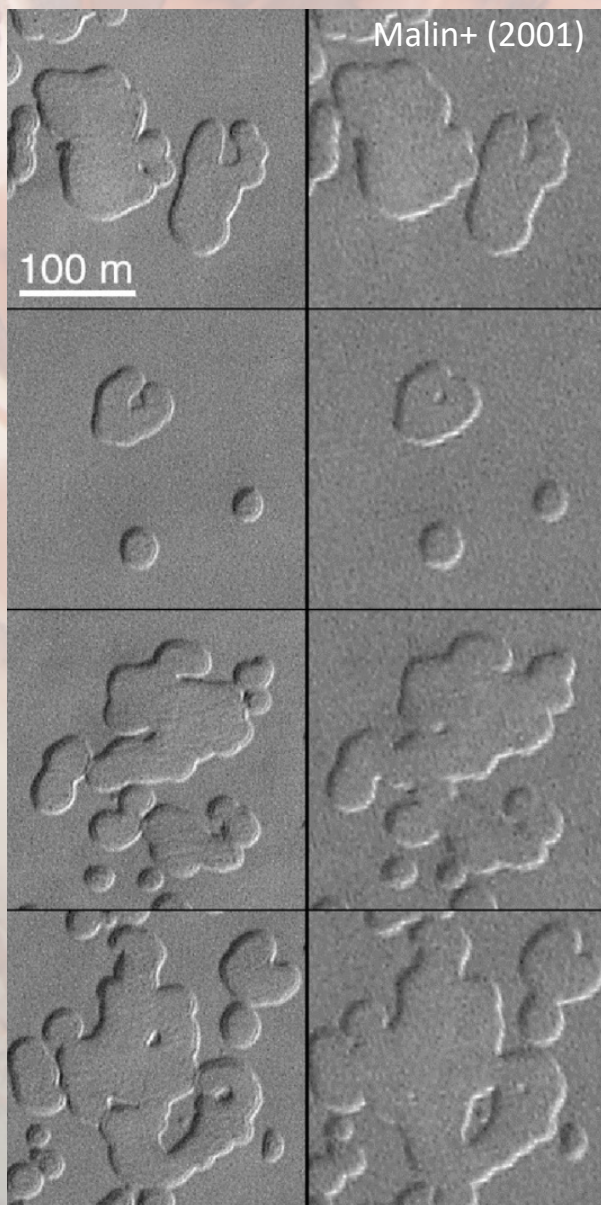
Mariner 9 Image Mosaic  
Murray & Malin, 1973



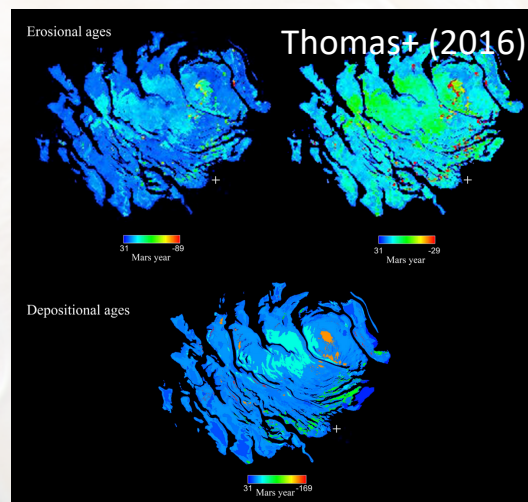
- “[From occultation] it can be seen that the southern residual cap must be higher than the northern one by at least 2 km.”
- “Any solid CO<sub>2</sub> in the south would be in contact with atmospheric CO<sub>2</sub> at a pressure lower by about 2 mbar than in the north.”
- “There is no reason to suppose a permanent CO<sub>2</sub> southern cap would be at a systematically lower temperature than the northern one.”
- “Hence, solid CO<sub>2</sub> deposits in the south would be out of equilibrium and would gradually be transferred to the north...in well under 1000 years.”
- “Excess solid carbon dioxide is probably present [buried] in the north residual cap.”



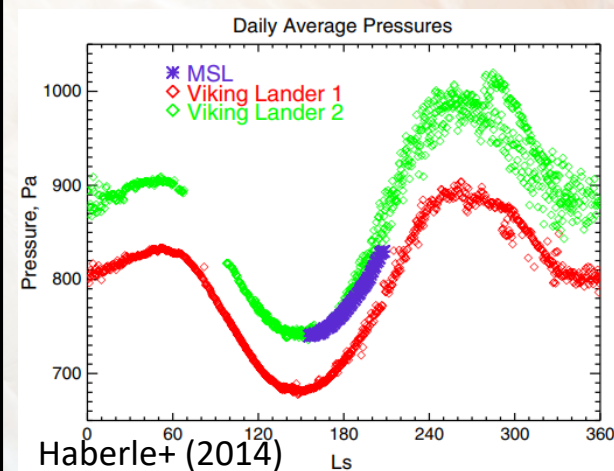
# Is the RSPC disappearing?



- “The erosion ... and other observations suggest that the present martian environment is neither stable nor typical of the past.” –Malin+ (2001)



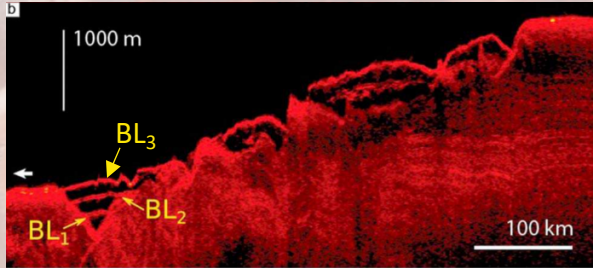
RSPC mass balance in  
Mars years 9–31  
~ -0.039% to +0.026%  
mean atmospheric CO<sub>2</sub>  
mass per martian year.



Viking vs. Curiosity pressure  
curves: no evidence yet for  
the 1–20 Pa increase  
expected from the possible  
loss of CO<sub>2</sub> from RSPC.

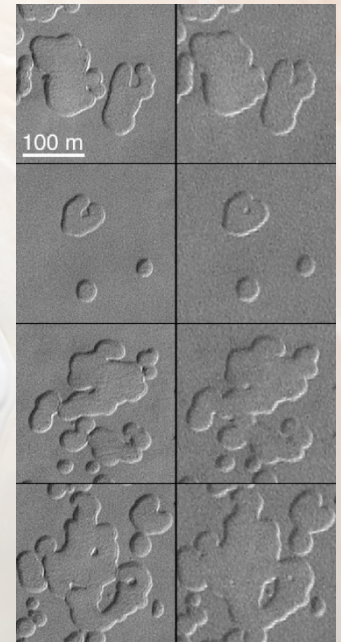


# 3 Outstanding Questions

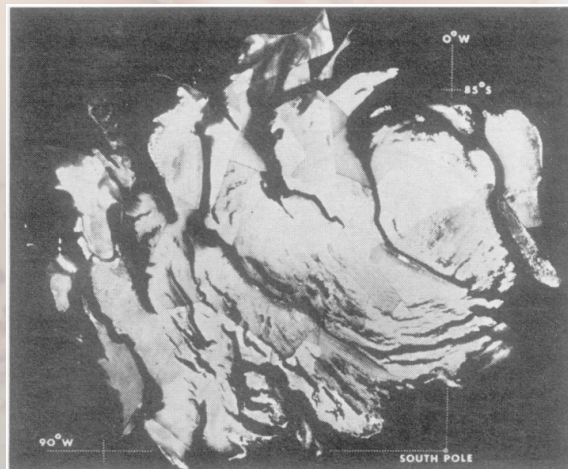


1. How was the massive CO<sub>2</sub> deposit emplaced with its observed stratigraphy?

2. Why does the RSPC exist?

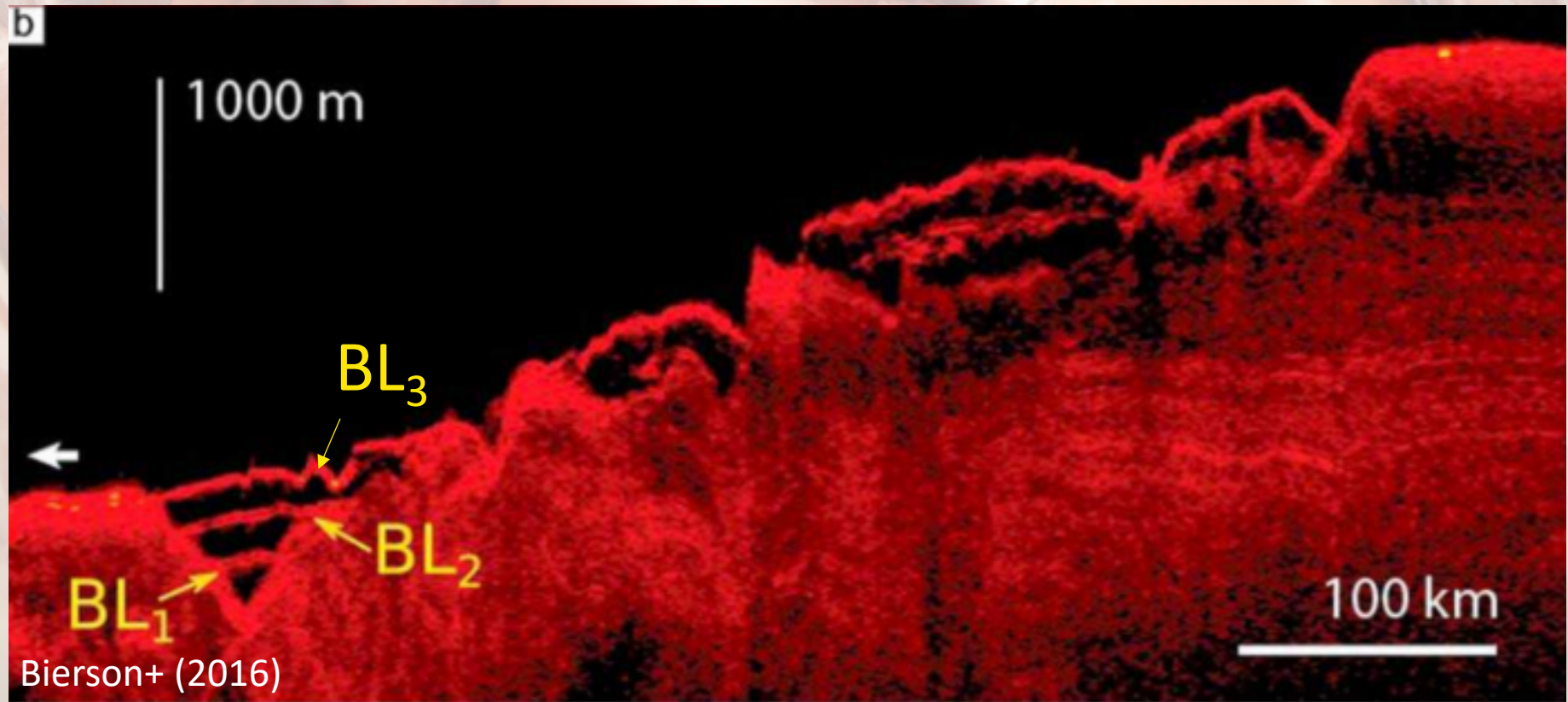


3. Why is permanent CO<sub>2</sub> at the south pole (not the north)?





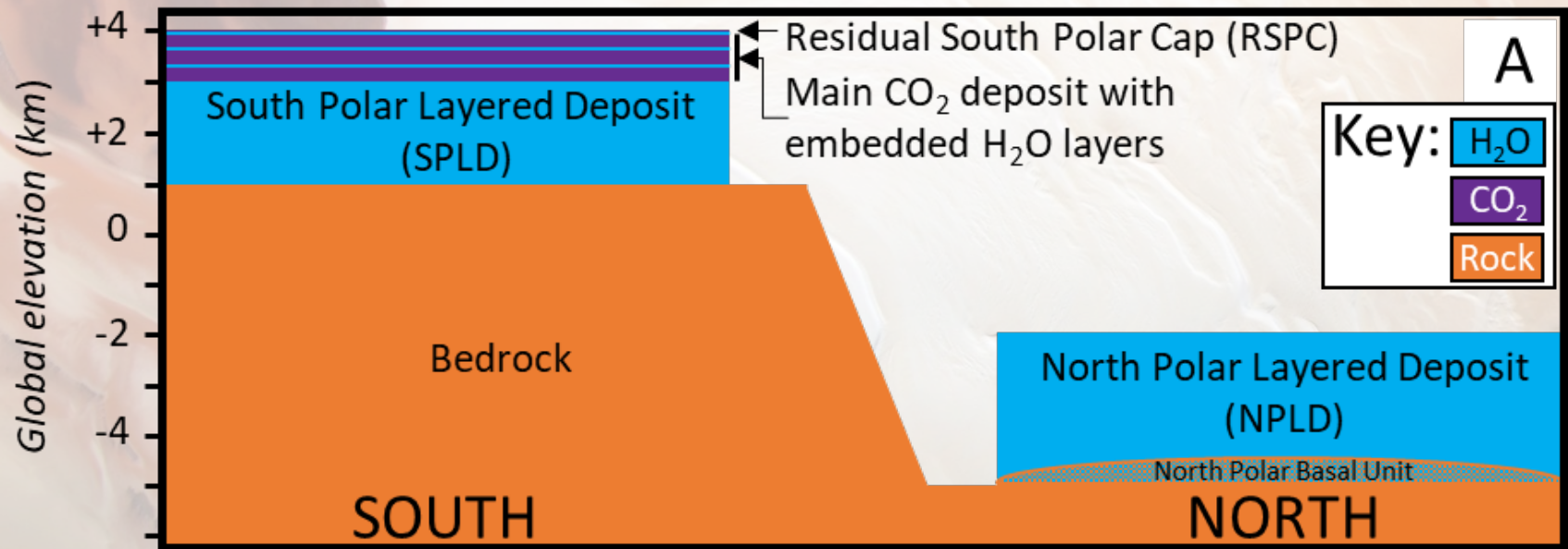
# The massive CO<sub>2</sub> deposit



“We find three distinct CO<sub>2</sub> subunits, each capped by a bounding layer (BL).” –Bierson et al. (2016)

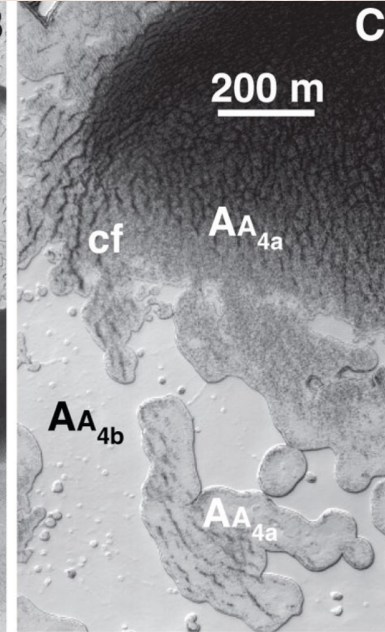
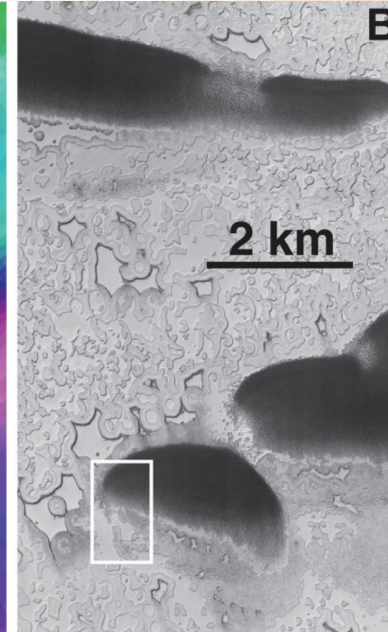
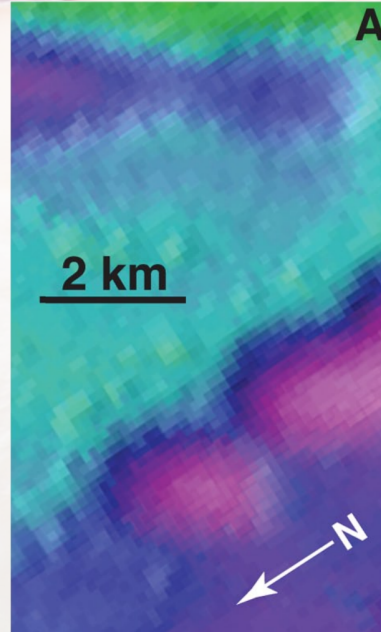
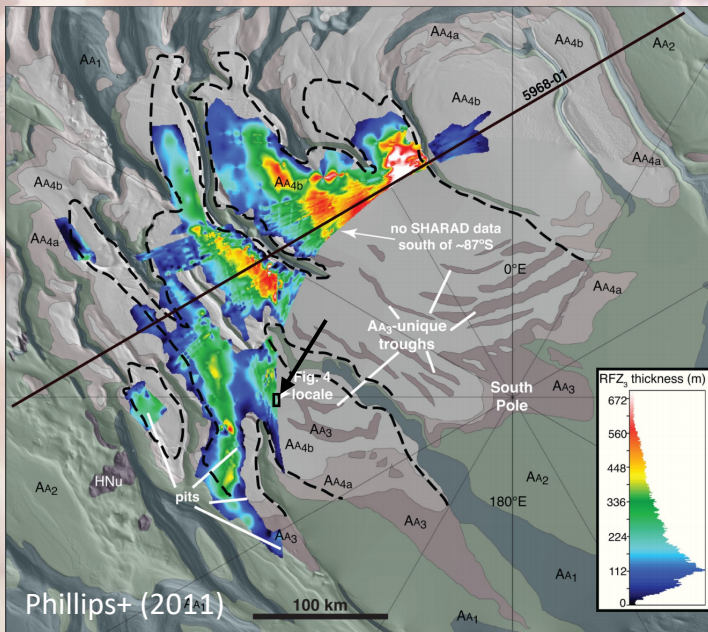
- “CO<sub>2</sub> ice is deposited over much of the poles during low obliquity periods.”
- “A remnant is sequestered below a water ice deposit (BL subunits), removing it from contact with the atmosphere.”

# Schematic Polar Stratigraphy

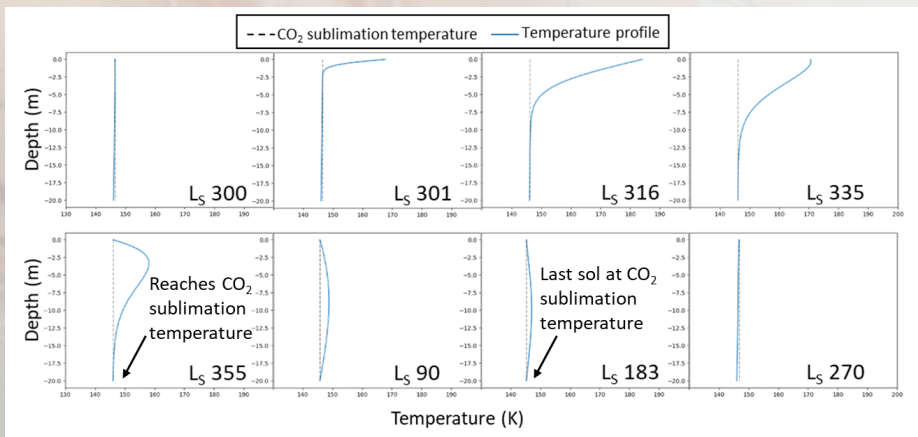




# Does the H<sub>2</sub>O seal the deposit?

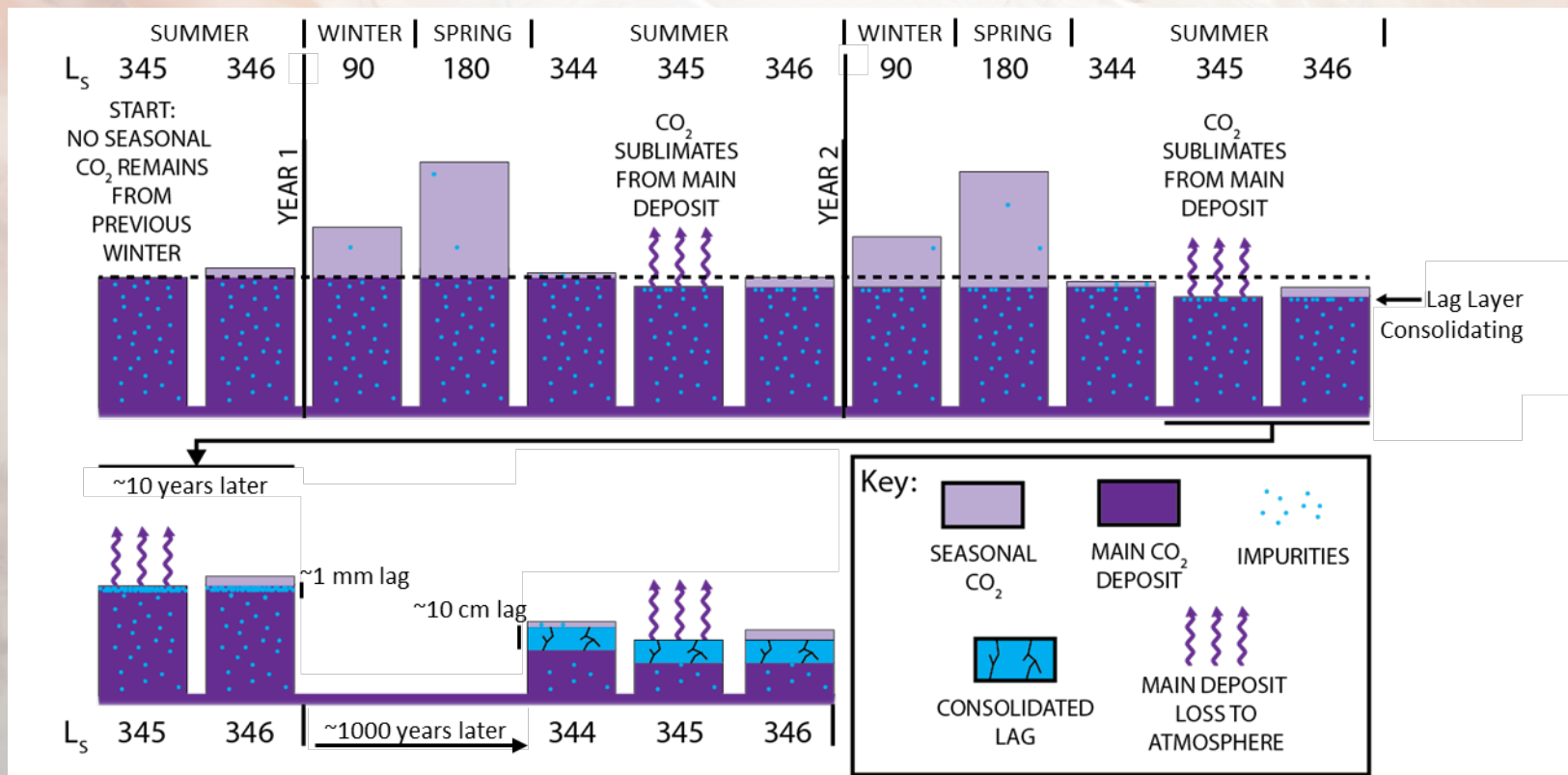


- “All of these smaller troughs, depressions, and pits appear to result from erosion and removal of unit A<sub>A3</sub> [the massive CO<sub>2</sub> deposit], with a strong component of sublimation and collapse.”
- “The fracturing, not found in other SPLD units, may be a response to continuing unit A<sub>A3</sub> sublimation after the pits had first formed.”



# Is the RSPC an extraordinary accident?

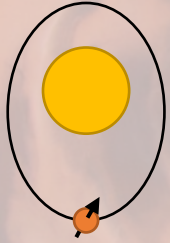
- If the massive CO<sub>2</sub> deposit is sequestered from the atmosphere, the same problem of “extraordinary accident” exists.
- *“A scum layer of dark (low albedo) material may be buried beneath a topmost layer of frost, but as soon as this topmost layer is removed, the dark dust [or H<sub>2</sub>O ice] will heat up and any CO<sub>2</sub> beneath it will escape.” –Murray and Malin (1973)*
- But what happens next?



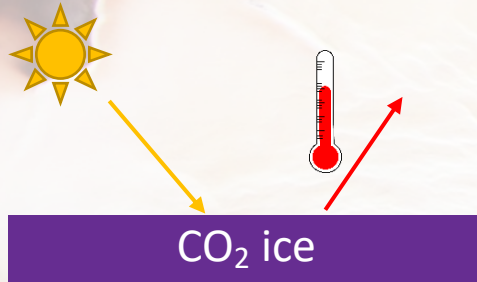


# Mars' Pressure History

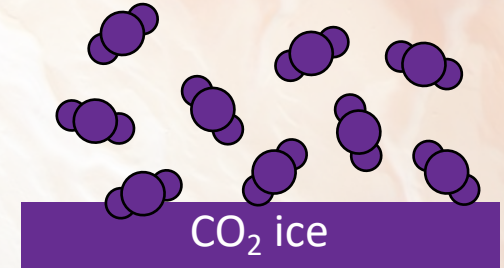
## Use a look-up table:



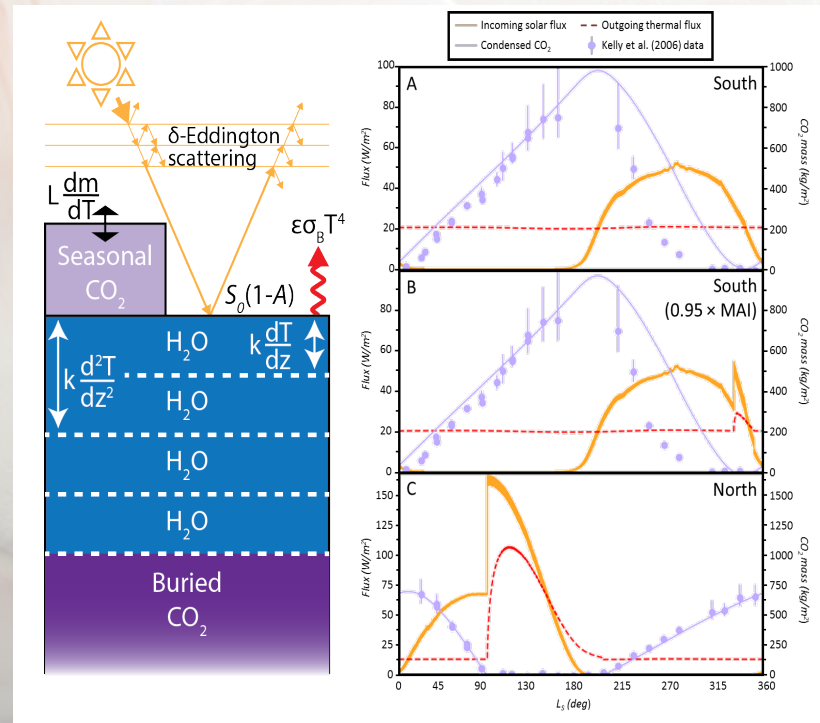
Any orbit



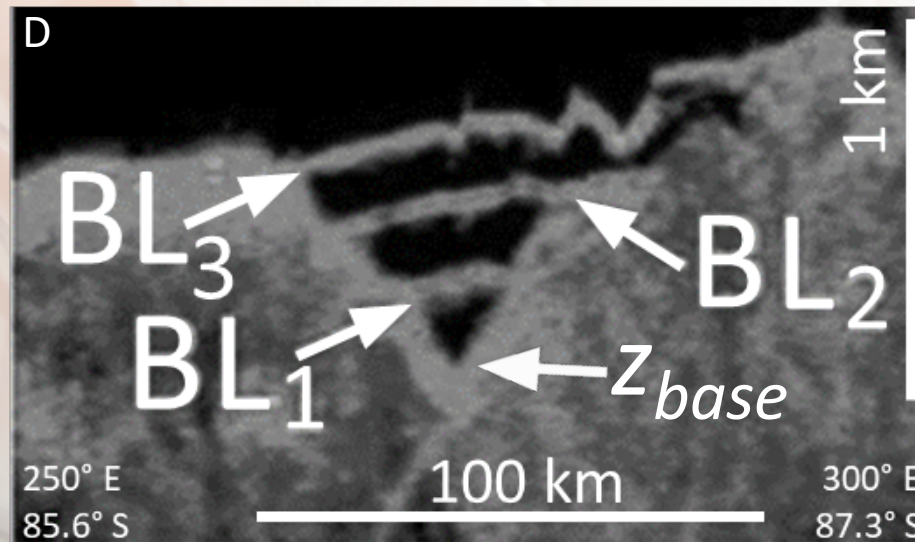
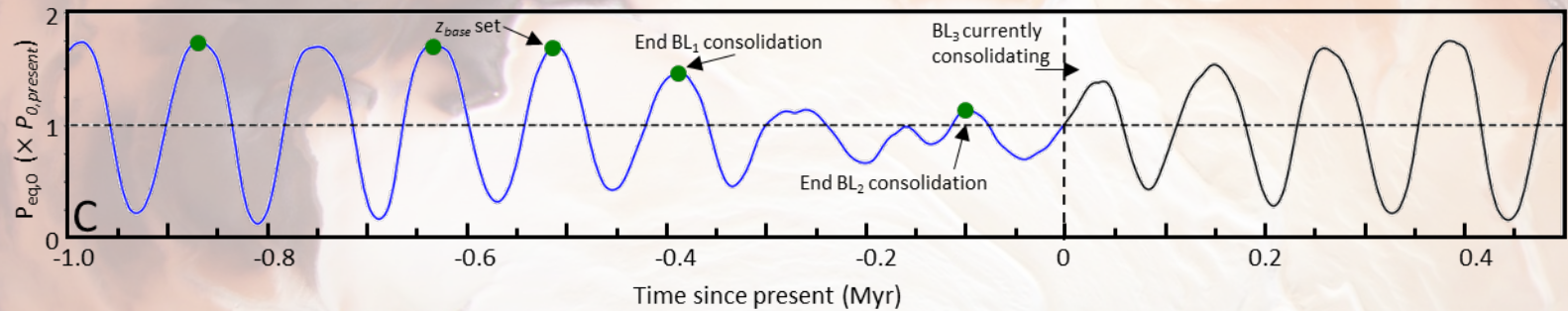
Equilibrium Frost  
Temperature



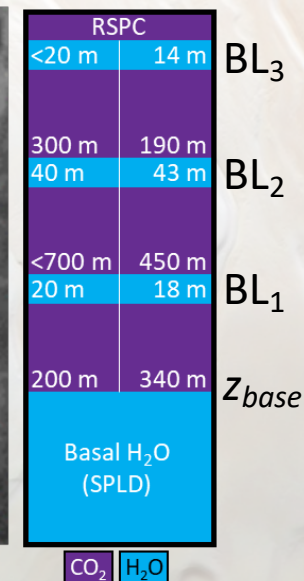
Equilibrium  
Pressure



# Mars' Pressure History and Stratigraphy

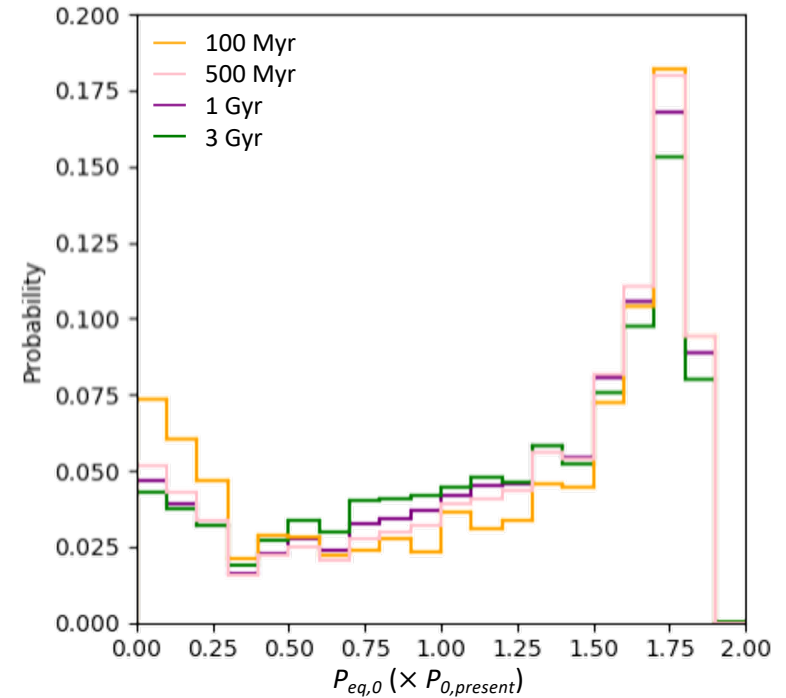
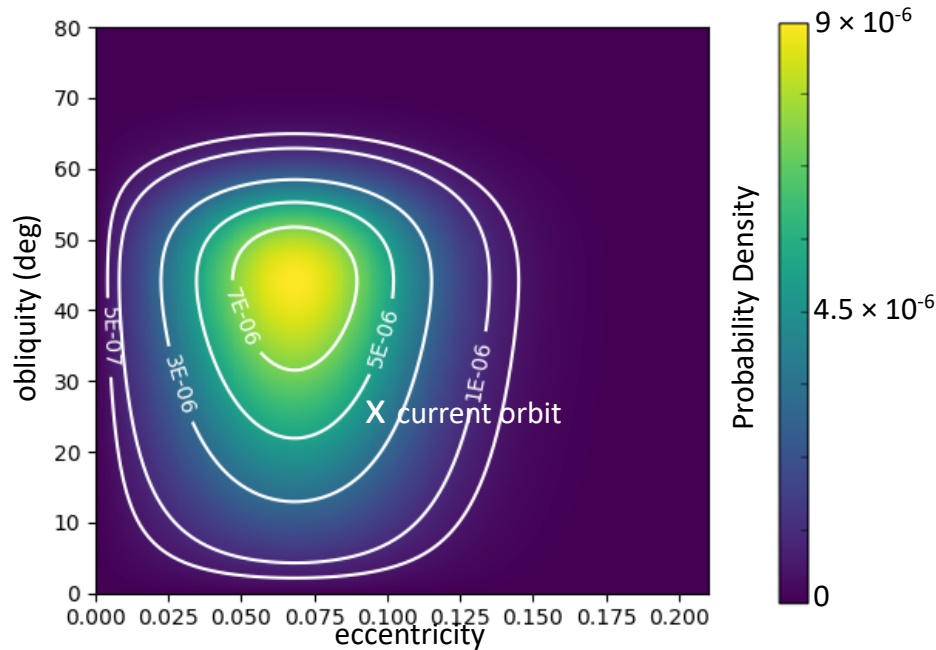


**E. Stratigraphy**  
Obs. Model





# Long-term pressure history of Mars



- Median Amazonian pressure:  $1.32 \times$  present
- Interquartile range: 0.77 to  $1.67 \times$  present

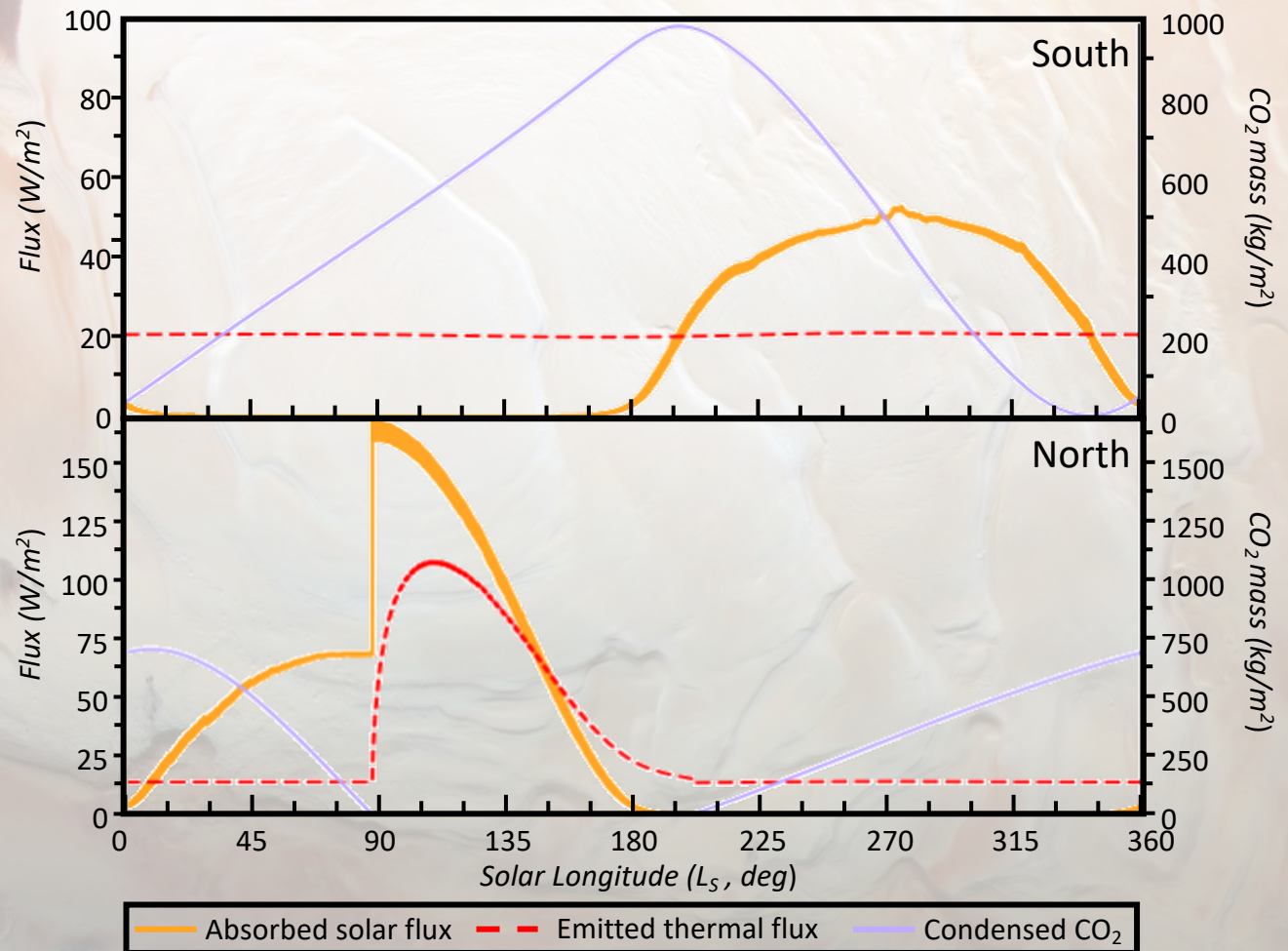
# Why is the permanent CO<sub>2</sub> in the south?

*"There is no reason to suppose a permanent CO<sub>2</sub> southern cap would be at a systematically lower temperature than the northern one."*

– Murray and Malin (1973)

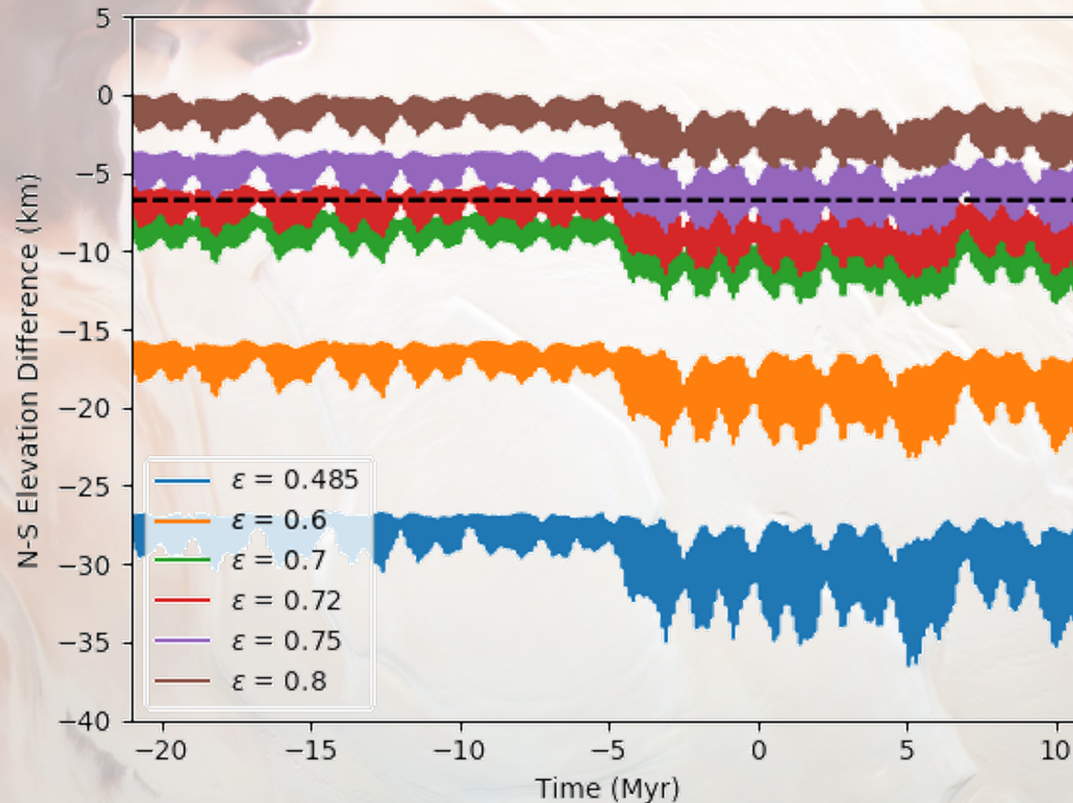
$$\bar{A}_{CO_2} \approx 0.7$$
$$\epsilon_{CO_2} \approx 0.8$$

$$\bar{A}_{CO_2} \approx 0.63$$
$$\epsilon_{CO_2} \approx 0.49$$



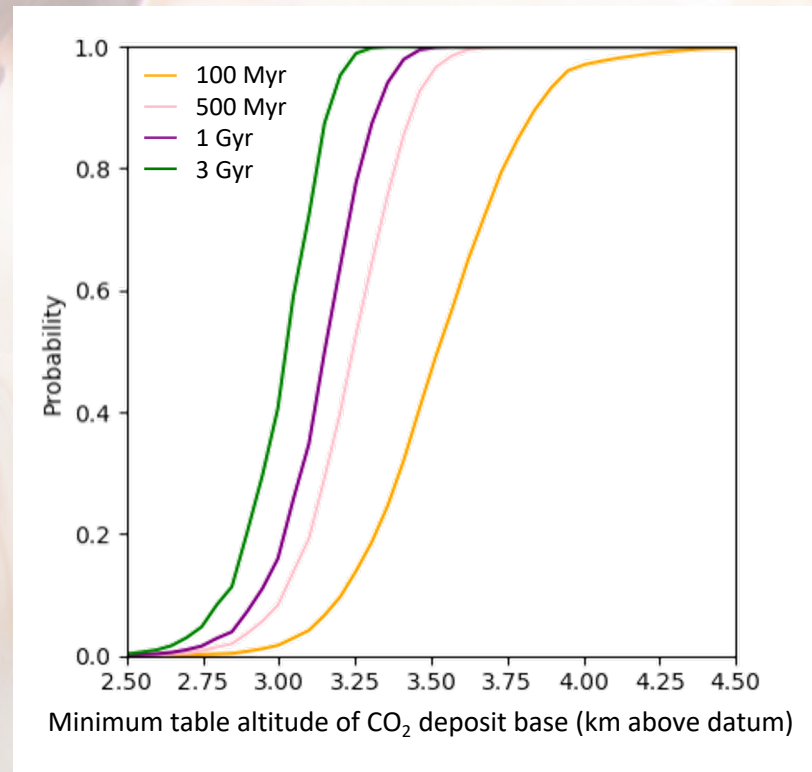


# Will perennial northern CO<sub>2</sub> ever exist?



Would require significant changes to  $\bar{A}_{CO_2}$  and/or  $\epsilon_{CO_2}$

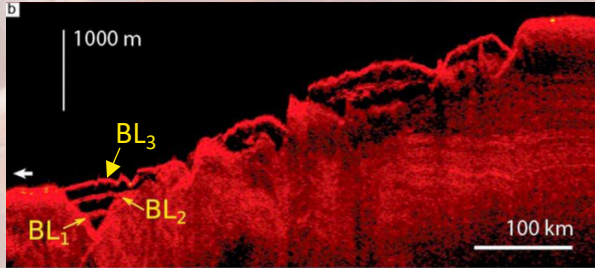
# CO<sub>2</sub> protects the SPLD at high obliquity



- CO<sub>2</sub> protects the SPLD from ablating.
- CO<sub>2</sub> does not protect the NPLD.



# Conclusions

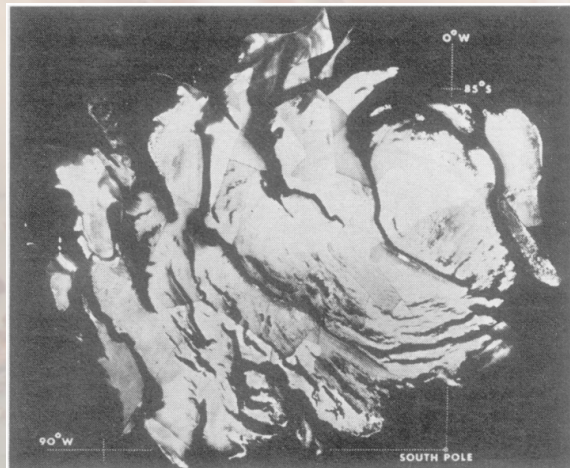
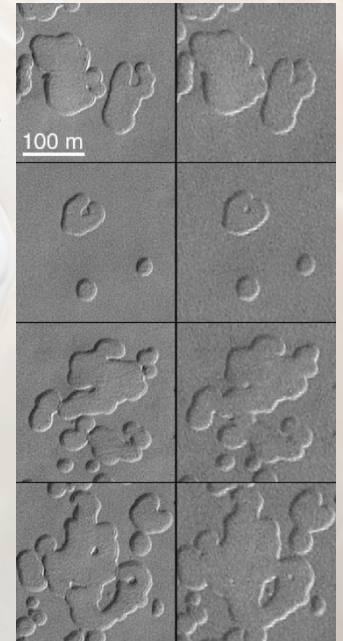


1. How was the massive CO<sub>2</sub> deposit emplaced with its observed stratigraphy?

**Equilibrated co-evolution with the atmosphere, driven by orbital forcing. H<sub>2</sub>O impurities accumulate into lag deposits.**

2. Why does the RSPC exist?

**Negative feedback between surface CO<sub>2</sub> ablation, dark lag formation, and basal CO<sub>2</sub> sublimation.**

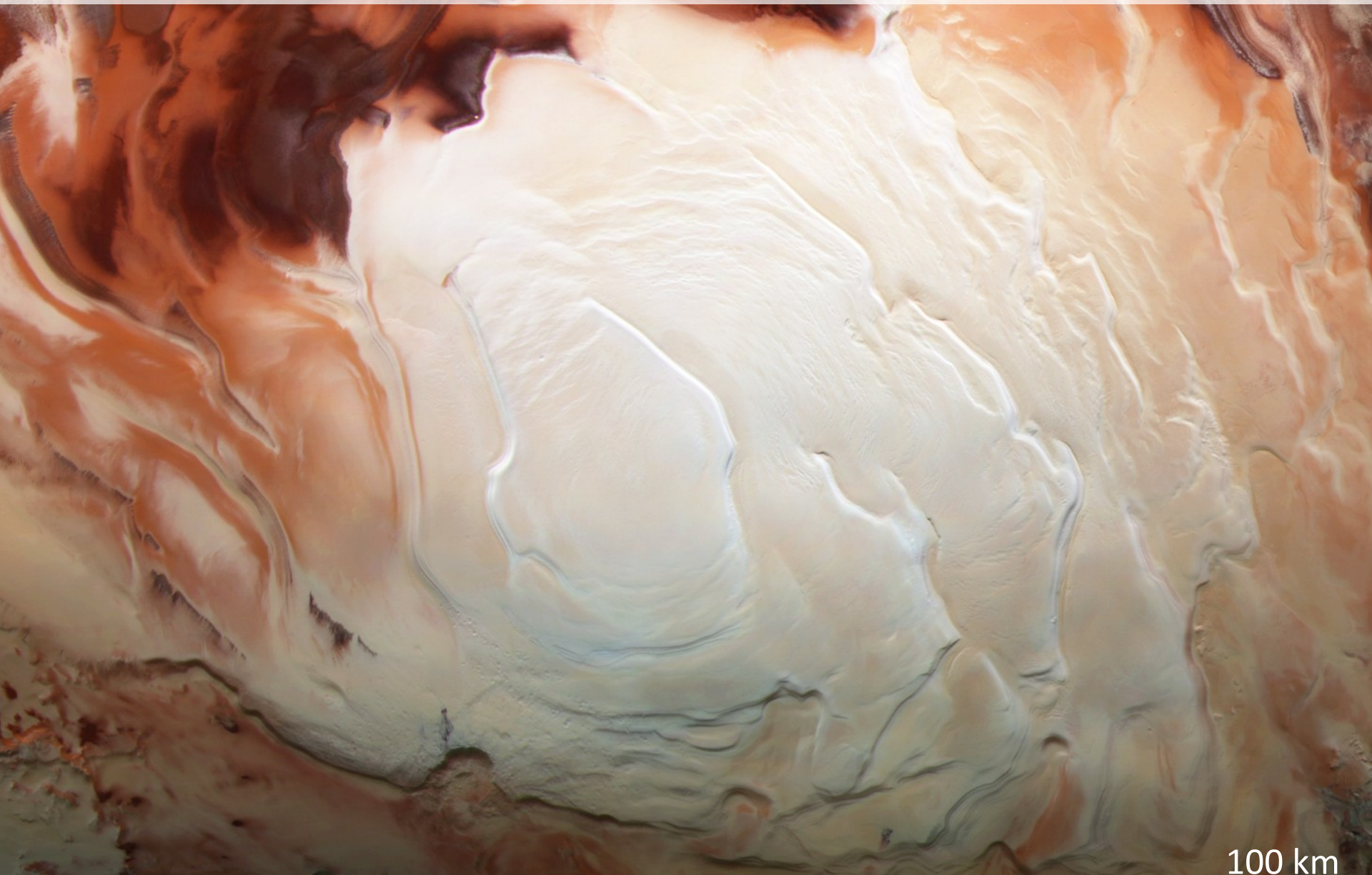


3. Why is permanent CO<sub>2</sub> at the south pole (not the north)?

**The albedo/emissivity of the southern CO<sub>2</sub> is higher, overwhelming the lower elevation of the northern cap.**



# Questions?



100 km

